

Brian Punsly

Position: Research Scientist Period covered: 12/2019-12/2020

I Scientific Work

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Black Holes and Quasars

1. Introduction

This report describes the research performed by Brian Punsly and collaborators in cooperation with ICRANet in 2020. The research was directed at finding environmental factors that are related to the switch-on of the general relativistic engine responsible for the few percent of accreting black holesthat drive powerful relativistic jets. This is important since this will relate directly to constraints on the initial state and boundary conditions on numerical models of black hole driven jets.

2. The Energetics of Launching the Most Powerful Jets in Quasars: A Study of 3C 82

Abstract:

 $3C \otimes 2 \otimes at \otimes a \otimes bift \otimes of \otimes 2.87 \otimes bite \otimes bite \otimes at \otimes 3C \otimes (Third \otimes Cambridge \otimes 2.87 \otimes bite \otimes 10^{-1})$ $Catalogue) \Box quasar. \Box Thus, \Box t \Box is \Box a \Box strong \Box candidate \Box to \Box have \Box the \Box most-luminous \Box$ $radio \Box lobes \Box in \Box the \Box universe. \Box 3C \Box 82 \Box belongs \Box to \Box the \Box class \Box of \Box compact \Box steep$ $spectrum \square radio \square sources. \square We \square use \square single-dish \square and \square interferometric \square radio \square observations \square$ $in \square order \square to \square model \square the \square plasma \square state \square of \square these \square powerful \square radio \square lobes. \square It \square is \square estimated \square the \square plasma \square state \square of \square these \square powerful \square radio \square lobes. \square It \square is \square estimated \square the 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\square plasma \square state \square plasma \square plasma \square state \square plasma \square plasma \square state \square plasma \square plasma \square state \square plasma plasma \square plasm$ that the long-term time-averaged jet power required to fill these lobes with leptonic plasma is $\overline{Q} \approx 2.66 \pm 1.33 \times 10^{47} \,\mathrm{erg \, s^{-1}}$, among the largest time-averaged jet powers \Box from \Box a \Box quasar. \Box Positing \Box protonic \Box lobes \Box is \Box not \Box tenable \Box as \Box they \Box would \Box $require \Box two \Box orders \Box of \Box magnitude \Box more \Box mass \Box transport \Box to \Box the \Box lobes \Box than \Box was \Box transport \Box to \Box the \Box transport \Box to \Box the \Box lobes \Box than \Box was \Box transport \Box to \Box the \Box lobes \Box than \Box was \Box transport \Box to \Box the \Box lobes \Box than \Box was \Box transport \Box to \Box the \Box lobes \Box than \Box was \Box transport \Box to \Box the \Box 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spectroscopic \square optical \square spectroscopic \square$ $is \square a \square powerful \square high-ionization \square broad-line \square wind \square with \square a \square kinetic \square power \square of \square ~10^{45} \square erg \square erg \square of \square ~10^{45} \square$ s^{-1} and a velocity of 0.01c. We also estimate from the broad lines in 2018 and \Box the \Box UV \Box continuum \Box in \Box three \Box epochs \Box spread \Box out \Box over \Box three \Box decades \Box that \Box the \Box accretion \Box flow \Box bolometric \Box luminosity \Box is $\Box L_{bol} \Box \approx \Box 3.2 \Box 5.8 \Box \Box \Box 10^{46} \Box erg \Box s^{-1}$. \Box The \Box ratio \Box $of \Box \overline{Q} / L_{bol} \approx 5.91 \pm 3.41 is \Box perhaps \Box the \Box largest \Box of \Box any \Box known \Box quasar. \Box Extremely \Box$ $powerful \exists jets \exists tend \exists to \exists strongly \exists suppress \exists powerful \exists winds \exists of \exists ionized \exists baryonic \exists matter. \exists$ Consequently, $\Box 3C \Box 82 \Box$ provides $\Box a \Box$ unique \Box laboratory \Box for \Box studying \Box the \Box dynamical \Box $\lim_{t \to 0} \int \frac{1}{2} \int \frac{$

3. The Extreme Red Excess in Blazar Ultraviolet Broad Emission Lines

ABSTRACT:

We present a study of quasars with very redward asymmetric (RA) ultraviolet (UV) broad emission lines (BELs). An excess of redshifted emission has been previously shown to occur in the BELs of radio-loud quasars and is most extreme in certain blazars. Paradoxically, blazars are objects that are characterized by a highly relativistic blueshifted outflow toward Earth. We show that the red emitting gas resides in a very broad component (VBC) that is typical of Population B quasars that are defined by a wide H β BEL profile. Empirically, we find that RA BEL blazars have both low Eddington rates (\gtrsim 1%) and an inordinately large (order unity) ratio of long-term time-averaged jet power to accretion luminosity. The latter circumstance has been previously shown to be associated with a depressed extreme UV ionizing continuum. Both properties conspire to produce a low flux of ionizing photons, two orders of magnitude less than typical Population B quasars. We use CLOUDY models to demonstrate that a weak ionizing flux is required for gas near the central black hole to be optimally ionized to radiate BELs with high efficiency (most quasars overionize nearby gas, resulting in low radiative efficiency). The large gravitational redshift and transverse Doppler shift result in a VBC that is redshifted by $\sim 2000-5000 \text{ km s}^{-1}$ with a correspondingly large line width. The RA BELs result from an enhanced efficiency (relative to typical Population B quasars) to produce a luminous, redshifted VBC near the central black hole.

2020 List of Publication

Punsly, Brian; Hill, Gary J.; Marziani, Paola; Kharb, Preeti; Berton, Marco; Crepaldi, Luca; Indahl, Briana L.; Zeimann, Greg, "The Energetics of Launching the Most Powerful Jets in Quasars: A Study of 3C 82",2020 ApJ 189 169

Punsly, B., Paola Marziani, Marco Berton, Preeti Kharb, "The Extreme Red Excess in Blazar Ultraviolet Broad Emission Lines",2020 ApJ903 44