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...start round table discussion ESOP XXXI...

Exposition tests with the new 535 nm filter

During ESOP XXVI in 2007 we made a workshop and decide to use standardized equipment for Beads observation. Every observer has to use:

- Optic 100/1000 Maksutov
- standardized filter made by the original green filter for 100/1000 + ND 4 layer
- video camera without automatic exposition
- exposition adaption by rotating crossed polarizer

IOTA/ES handle the order for optics and filter production and 2008 August 1 we place 3 observers on each limb in Sibiria for the best observed solar eclipse ever.

The standardized equipment idea was communicated published amongs others in the "Beads atlas".



100/1000 + watec piggypack on my telescope

Old equipments:

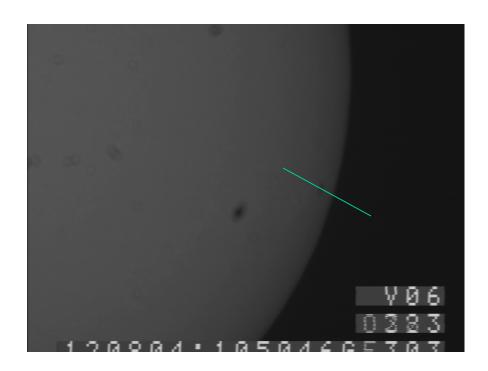
- 100/1000 + IOTA/ES filter
- Watec 120N $\gamma \rightarrow$ off, gain = 0
- no polarizer → sun in saturation and no sunspots visible

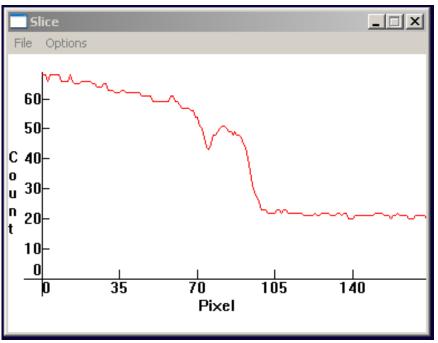
So do we **need crossed polarizer** to adapt the exposition.

During ESOP XXX we have got the contribution from A. Raponi. To unify our pictures with the Picard experiment we decide to install a small bandfilter 535 nm (now called "Raponi-filter").

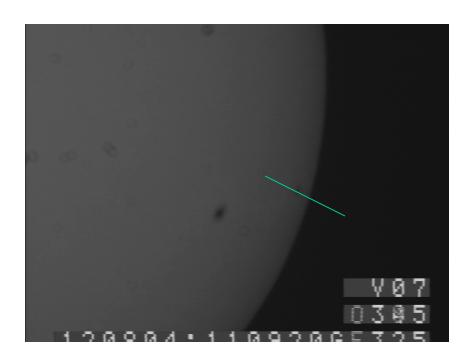
Result: not enoungh light with the old set up.

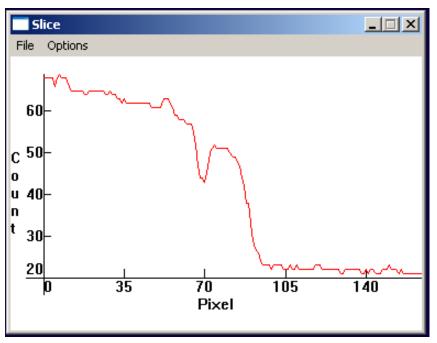
- 100/1000 + IOTA/ES filter + 535 nm
- Watec 120N $\gamma \rightarrow$ off, gain = 0
- polarizer parallel



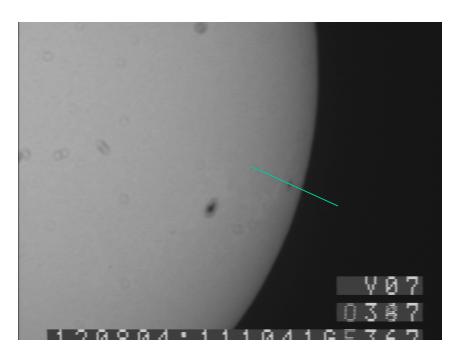


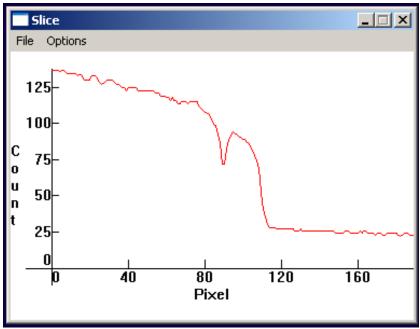
- 100/1000 + IOTA/ES filter + 535 nm
- Watec 120N $\gamma \rightarrow$ off, gain = 0.2
- polarizer parallel



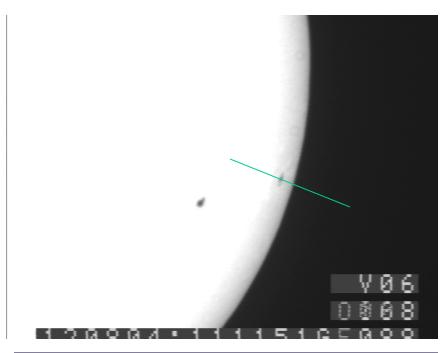


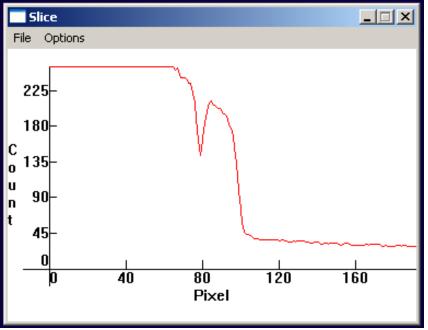
- 100/1000 + IOTA/ES filter + 535 nm
- Watec 120N $\gamma \rightarrow$ off, gain = 0.3
- polarizer parallel



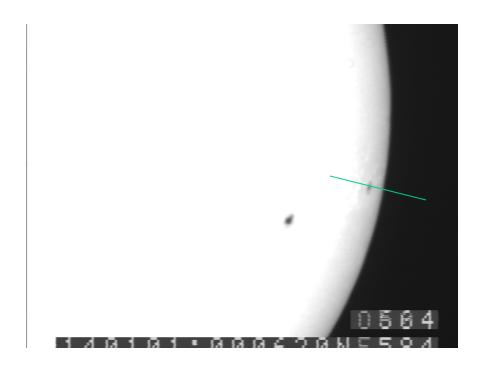


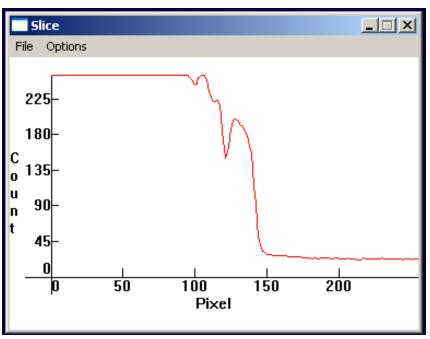
- 100/1000 + IOTA/ES filter + 535 nm
- Watec 120N $\gamma \rightarrow$ off, gain = 0.4
- polarizer parallel

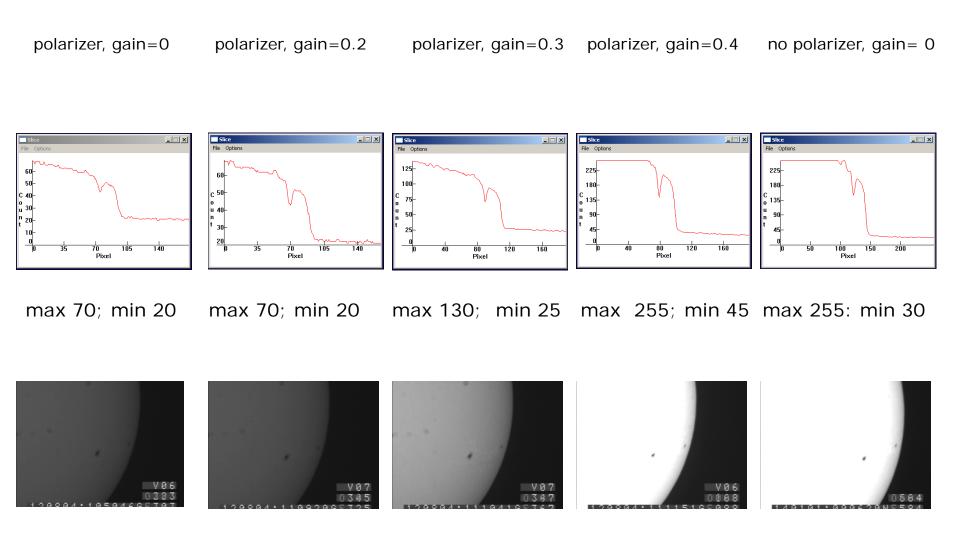




- 100/1000 + IOTA/ES filter + 535 nm
- Watec 120N $\gamma \rightarrow$ off, gain = 0
- no polarizer







Center of solar disk shall be in saturation \rightarrow remove polarizer or turn on gain.